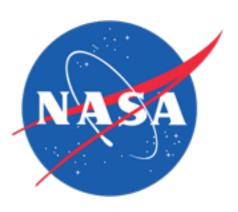
The dynamic nature of our Sun

Dr. Laurel Rachmeler, NASA/MSFC Osher Lifelong Learning Institute Our Place in the Universe 4 February 2016

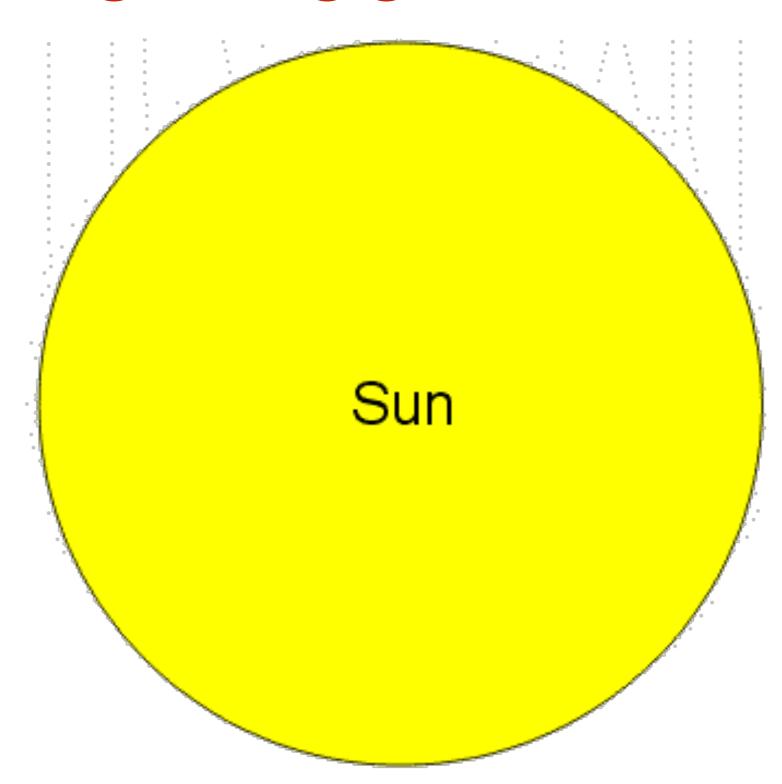


Quiz

How many Earths across is the Sun?

10	100	1000

Answer: 100





Quiz

How long does it take the Sun's light to reach the earth?

0.008 sec	8 sec	8 min

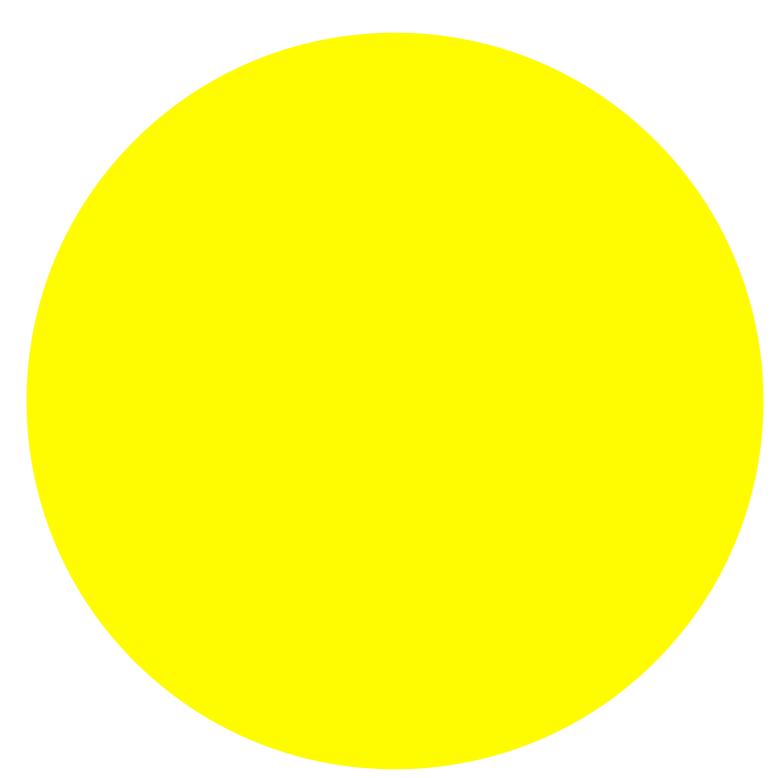
Answer: 8 minutes

```
distance to the Sun = 1 astronomical unit
= 93 million miles
```

= 150 million km

Light travels 300,000 km/s

Our Sun



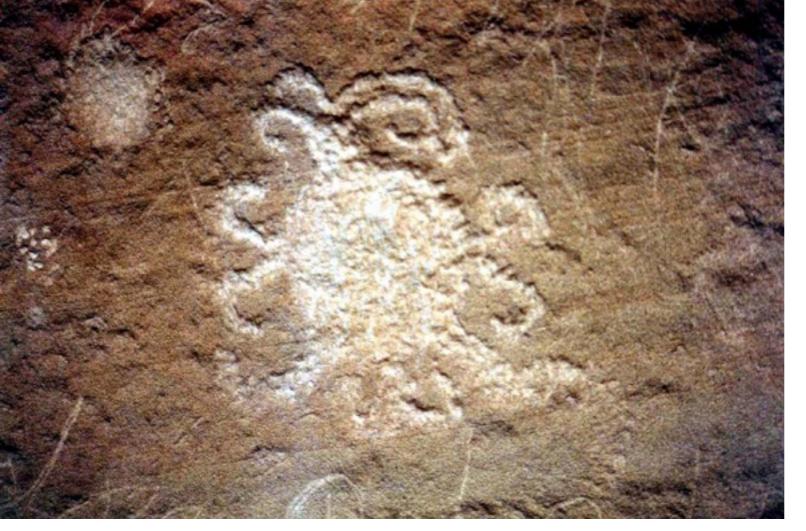


Solar Eclipse

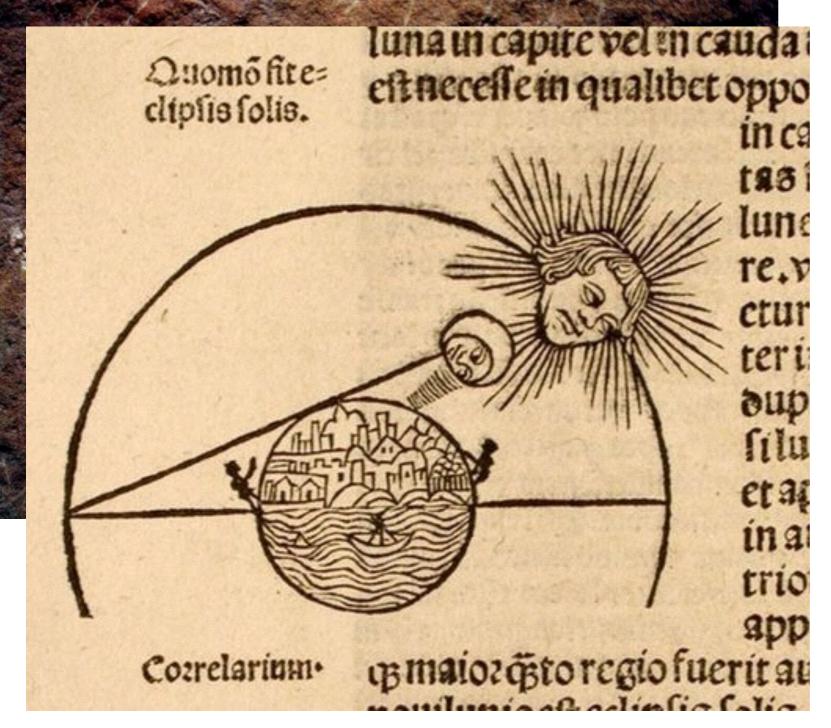




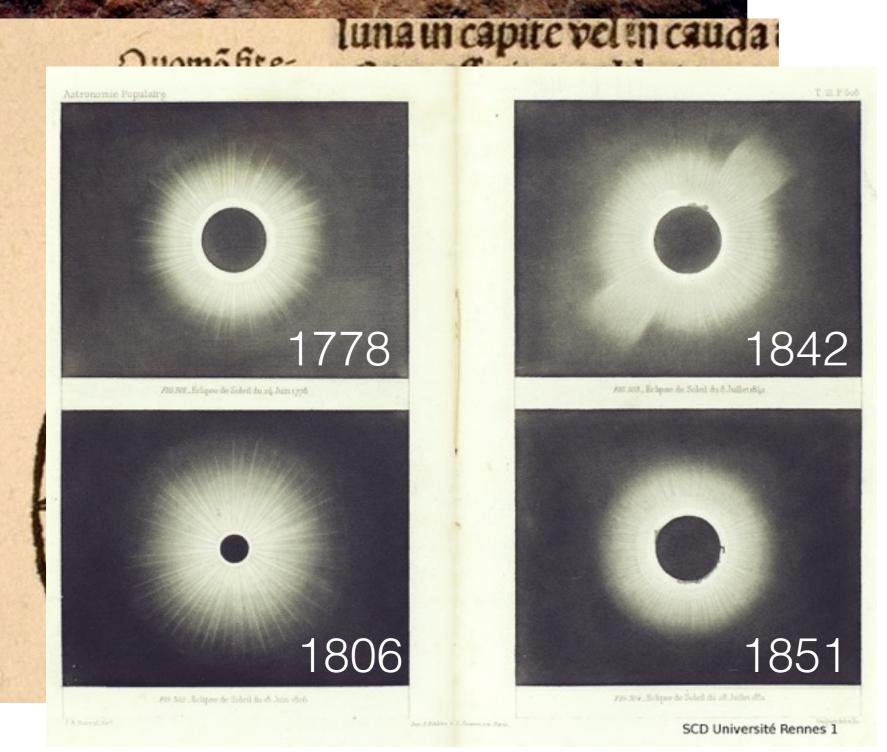
Moon



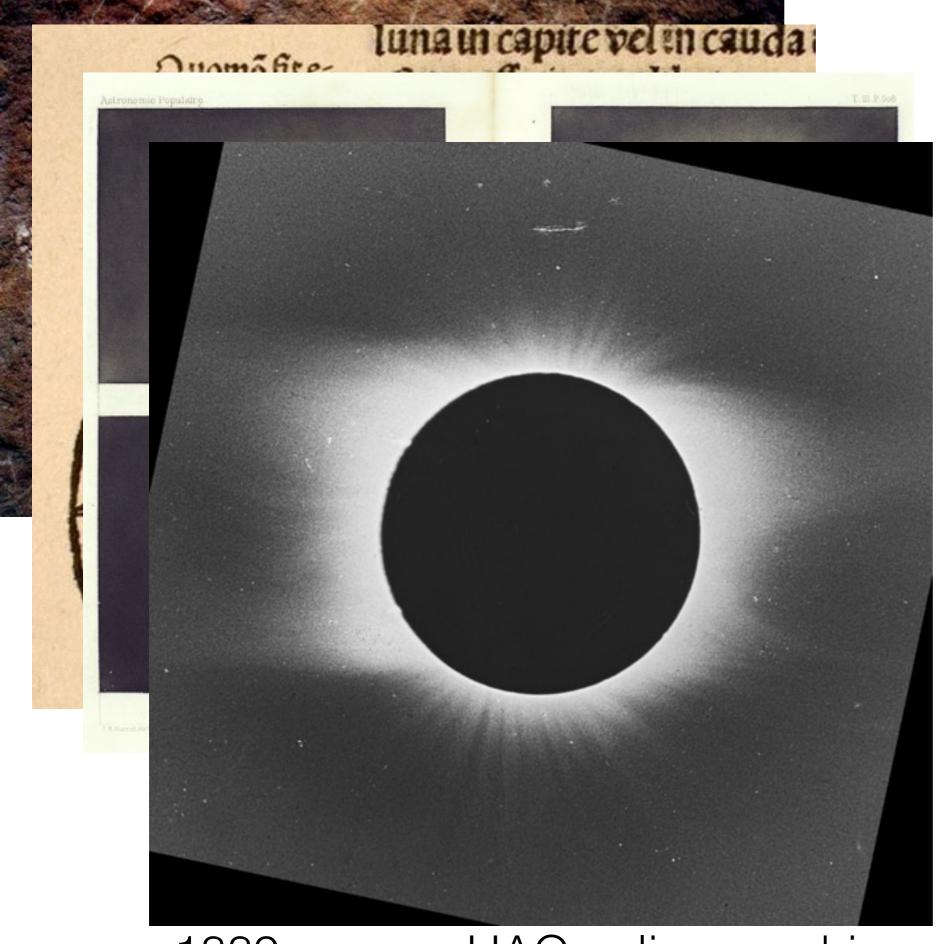
Petroglyph ~1000 AD (source HAO)



1499, Sacro Busto



Popular Astronomy



1889 source: HAO eclipse archive



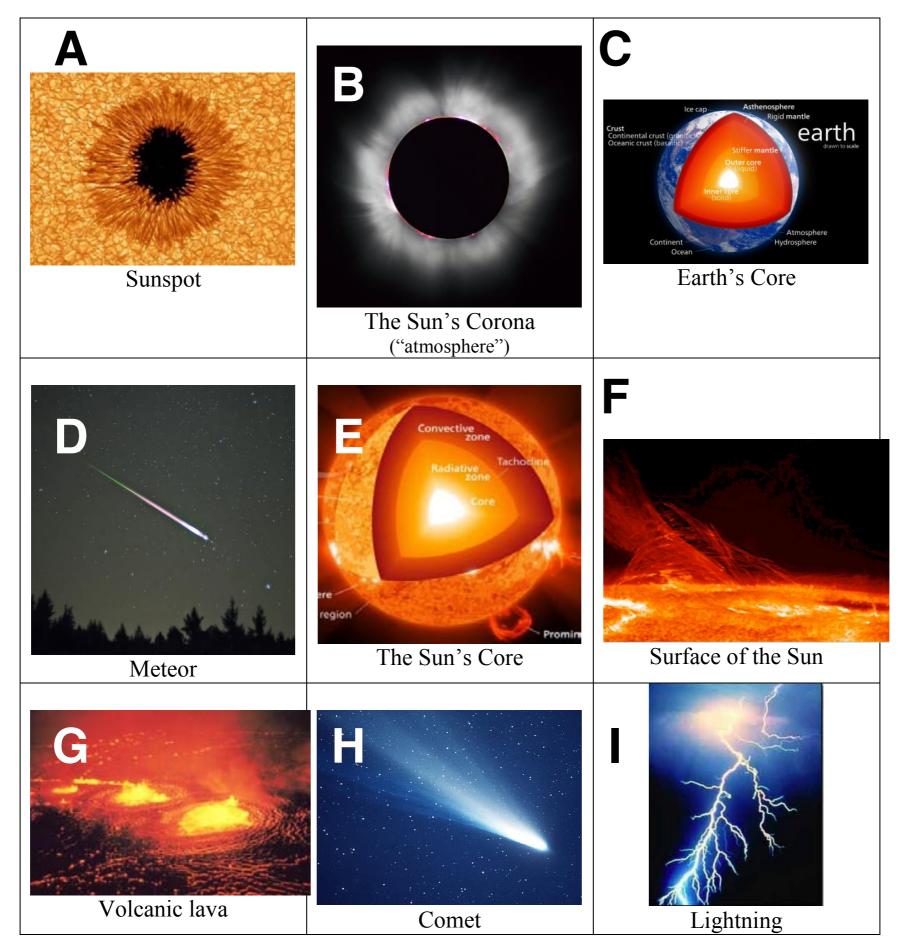


Activity

Get into groups of 4 or 5.

Record your answers.

Order from coolest to hottest



Answer

Comet -450°F to 200°F

Crust
Continental crust (principle)
Continental crust (principle)
Suifier mantle
Outer core

Earth's core 6200°K

G



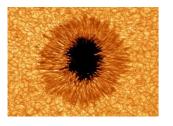
Lava 1450°F to 2000°F

I



Lightning 30,000°K

A



Sunspot 6300°F

В



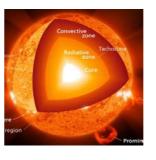
Sun's corona 5 million °K

D

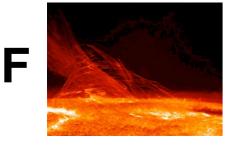


Meteor 10,000°F or 5800°K

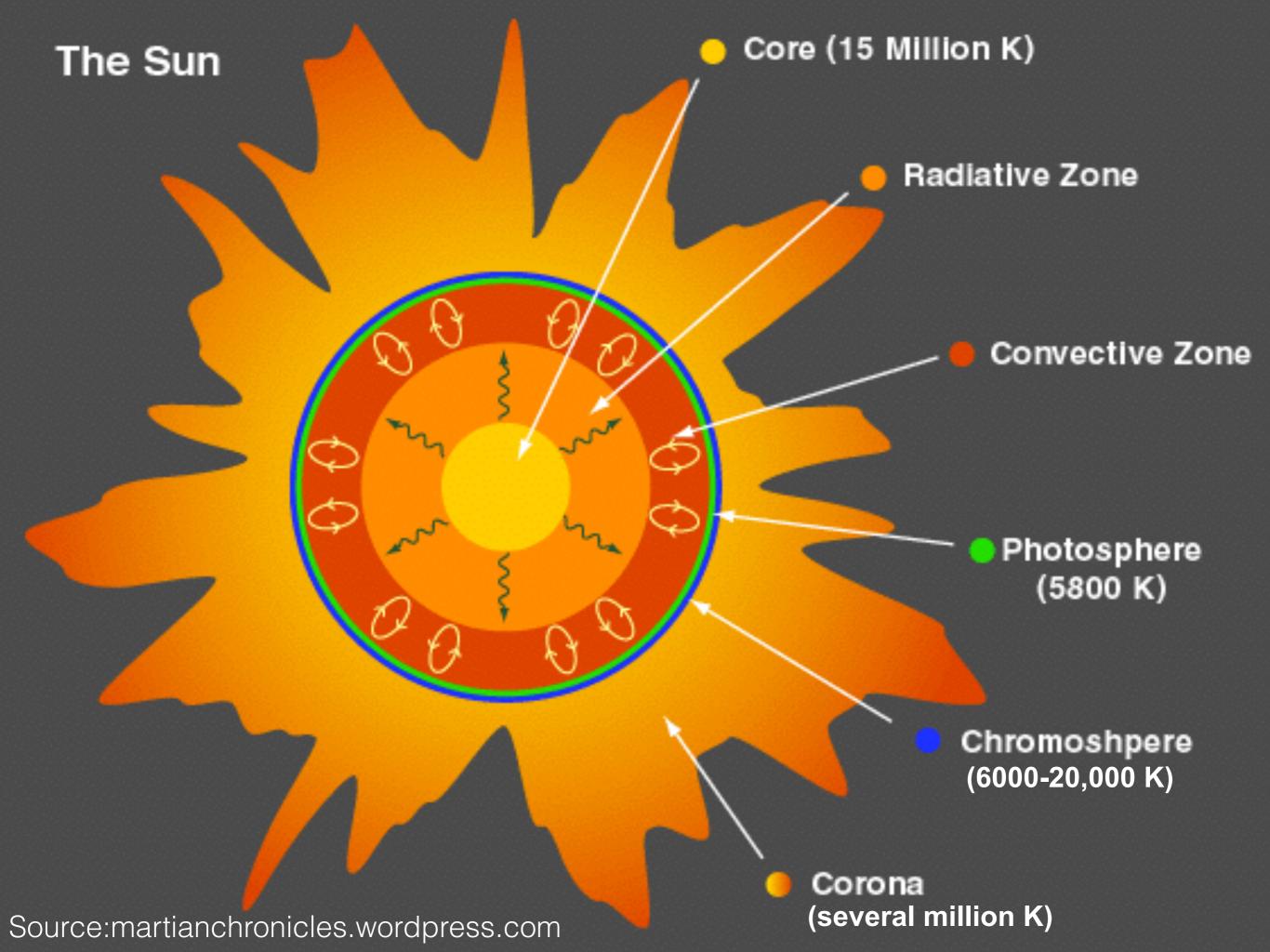
Ε



Sun's core 15 million °K



Sun's surface 6000°K

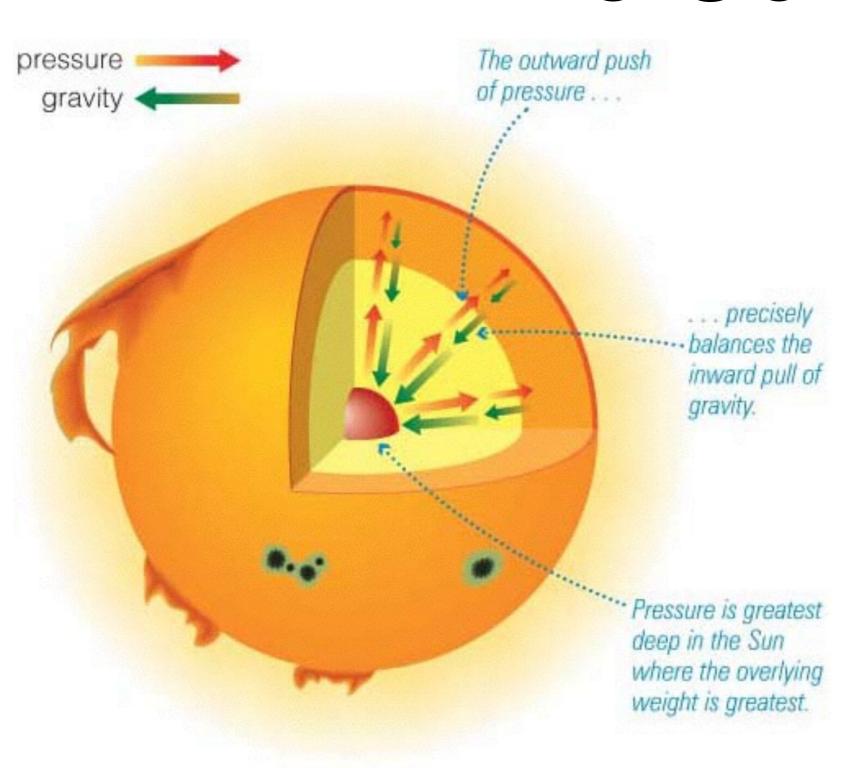


Plasma

source: UC Regents

PHASES OF MATTER proton (+ charge) nucleus electron neutron (no charge) structure of an atom Solid Liquid Gas Plasma LOW HIGH **Temperature or Energy**

The Core

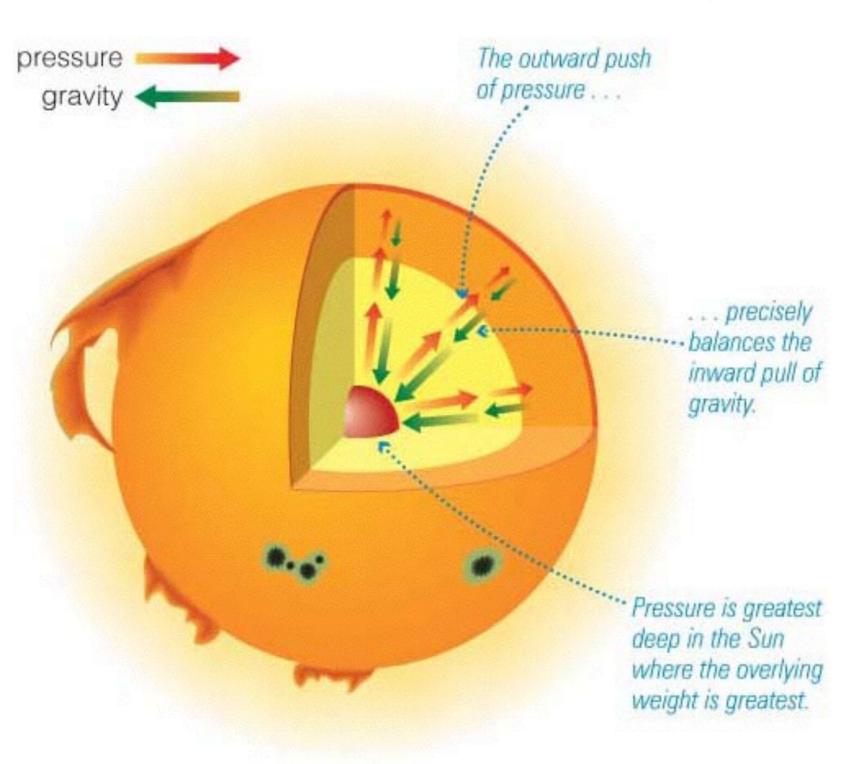


Thermal pressure balances gravity.

Pressure, and density, are greatest at the core.

source: http://lasp.colorado.edu

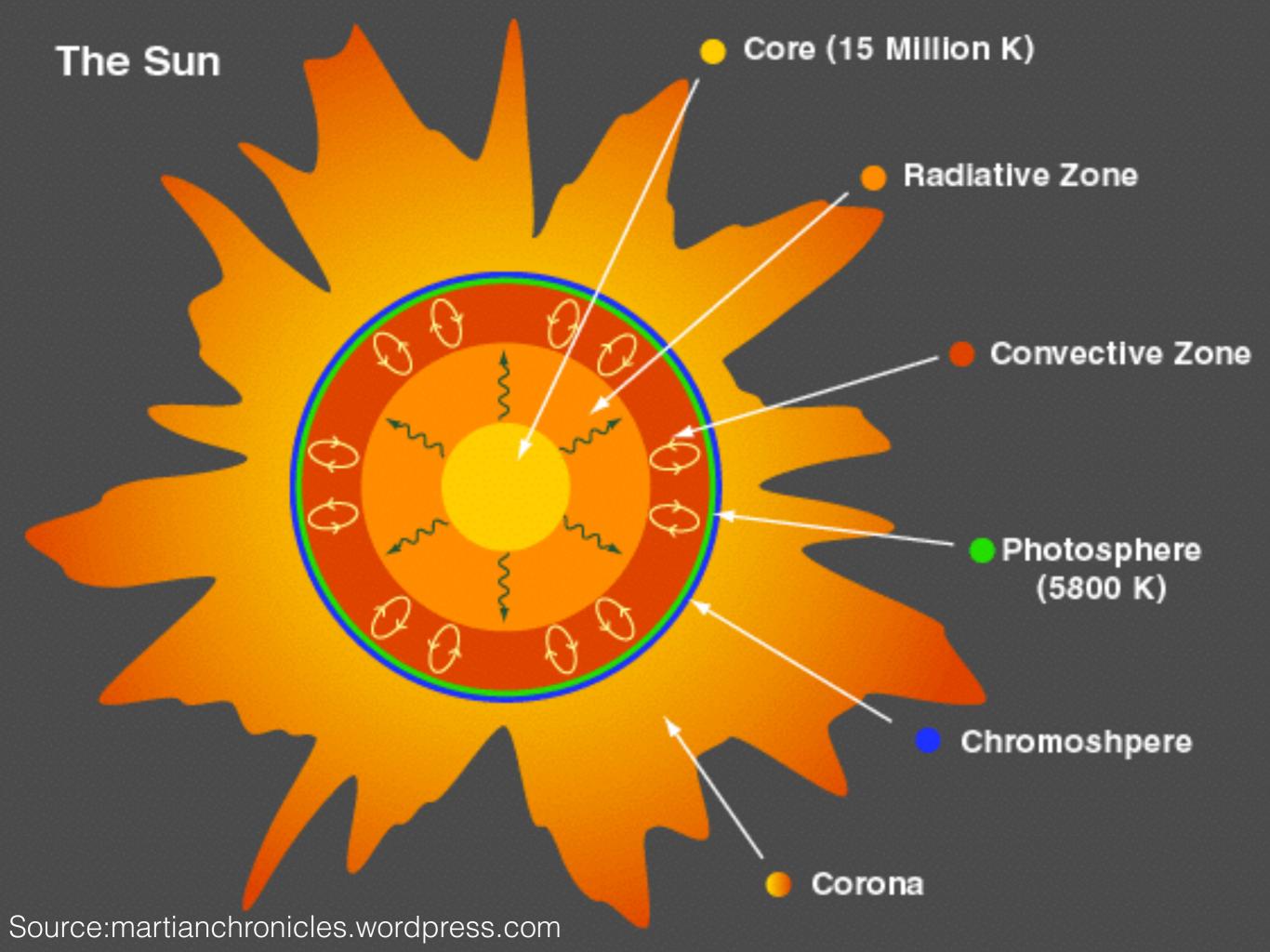
The Core



High densities at the core enable nuclear *fusion*.

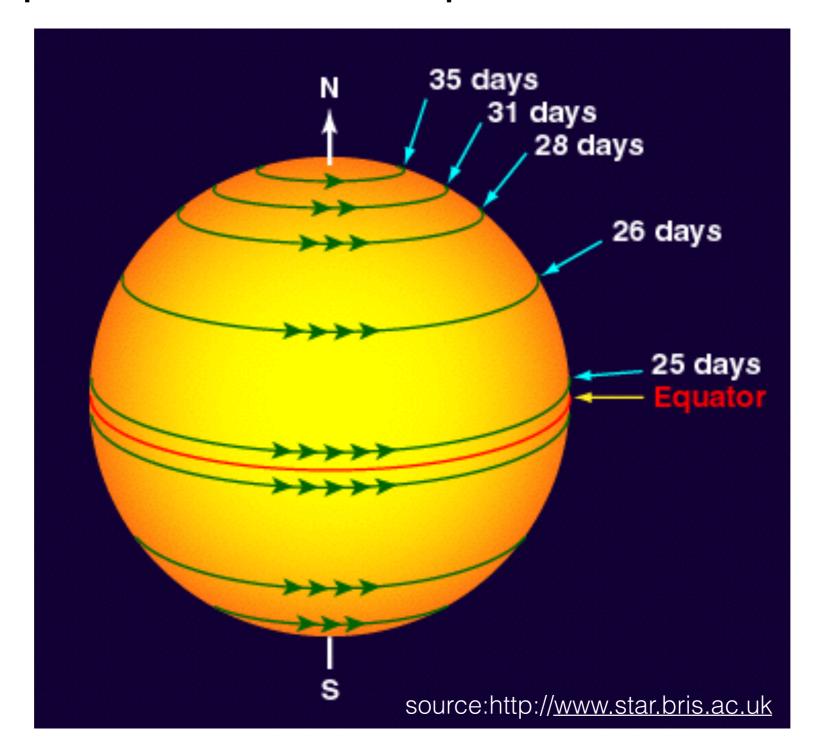
Hydrogen is combined into Helium, releasing energy (E=mc²) as light and heat.

source: http://lasp.colorado.edu



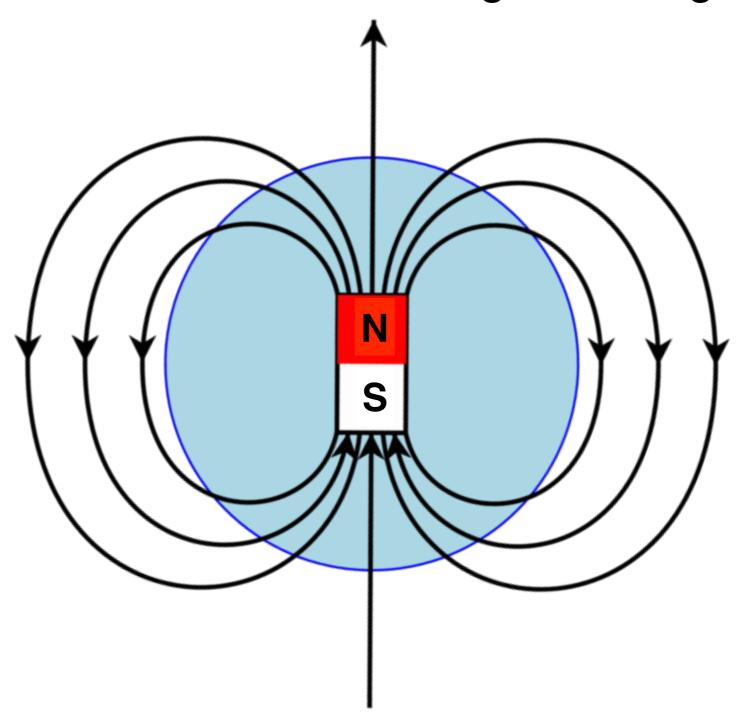
Solar Rotation

The equator spins faster than the poles: differential rotation.



Magnetic field

The Sun, like the Earth, has a global magnetic field.





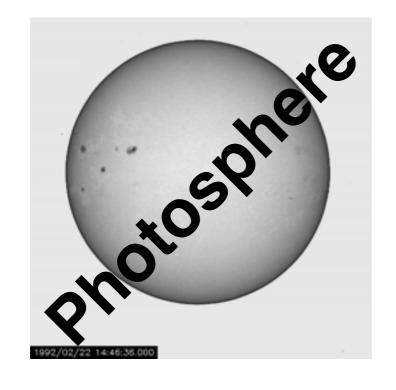
Solar Dynamo

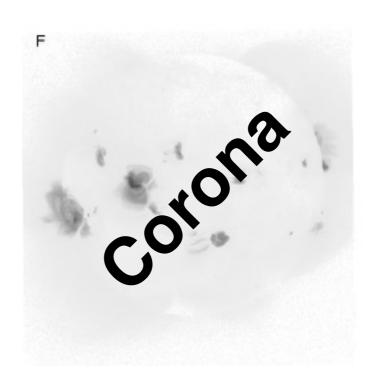
Spinning plasma drags the magnetic field.

Activity

Get into groups of 4 or 5.

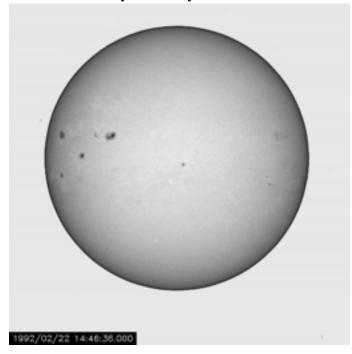
Match each photosphere picture to the corresponding corona picture.



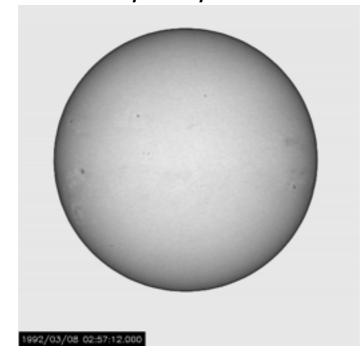


Answer

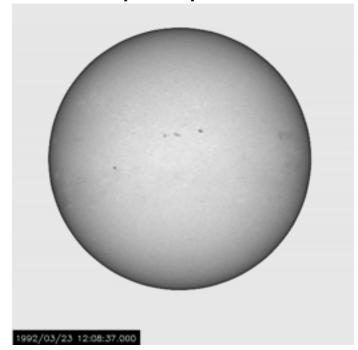
1992/02/22



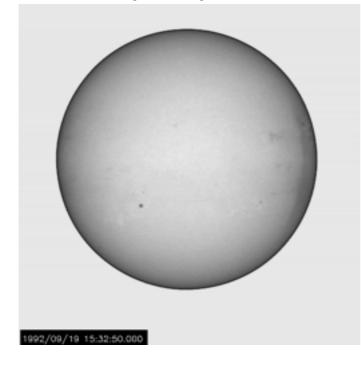
1992/03/08



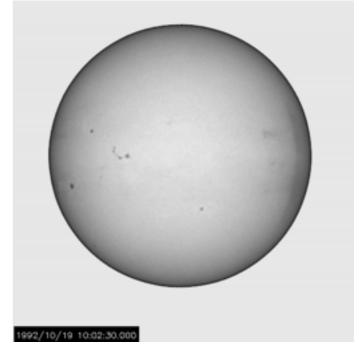
1992/03/23



1992/09/19

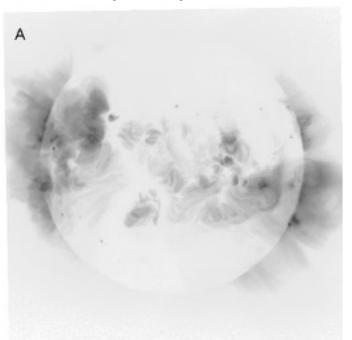


1992/10/19

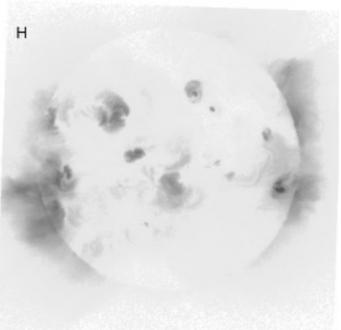


Answer

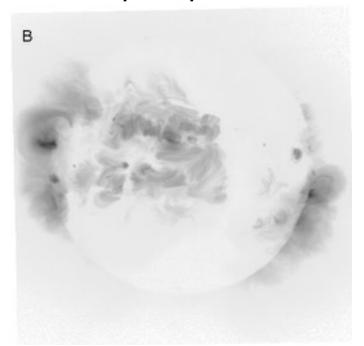
1992/02/22 **A**



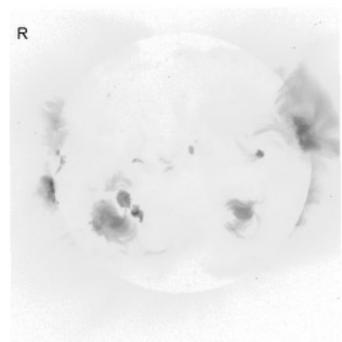
1992/03/08 **H**



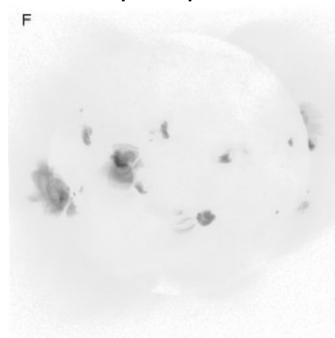
1992/03/23 **B**



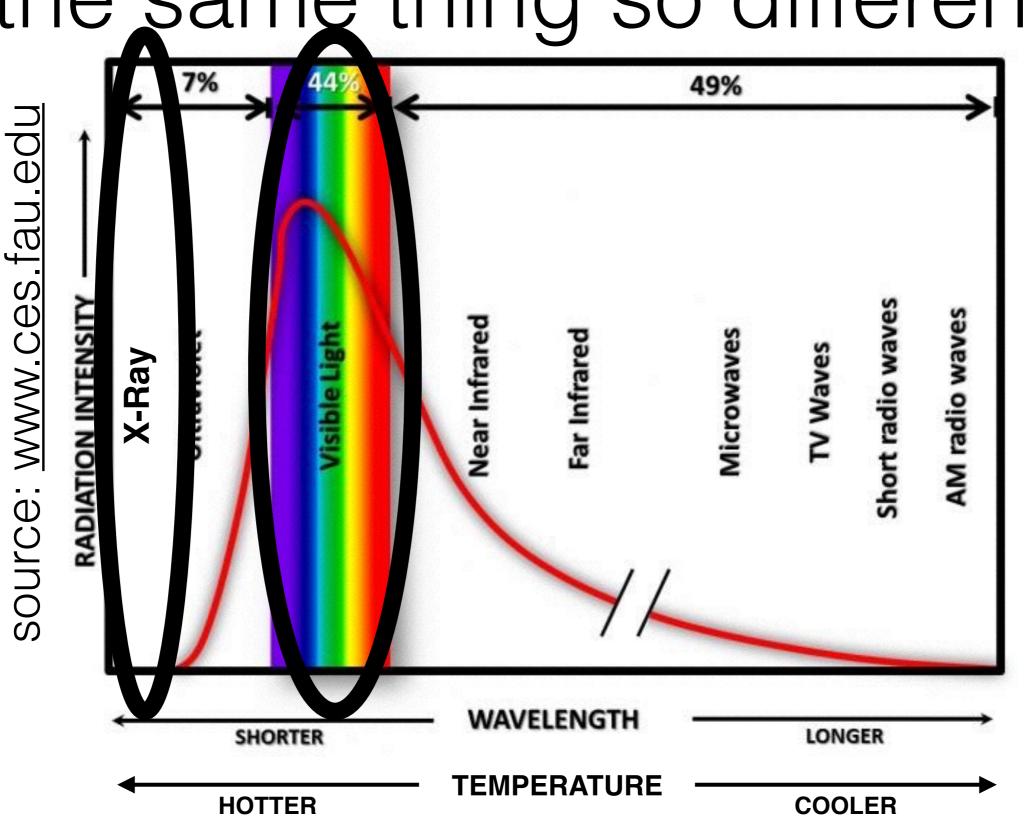
1992/09/19 **R**

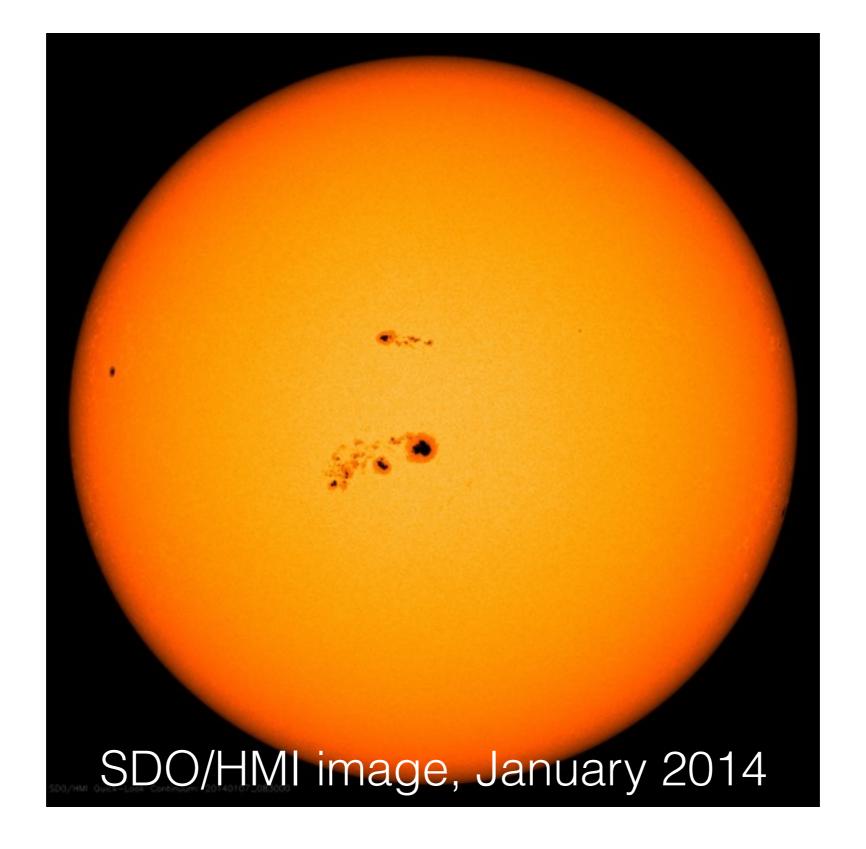


1992/10/19 **F**

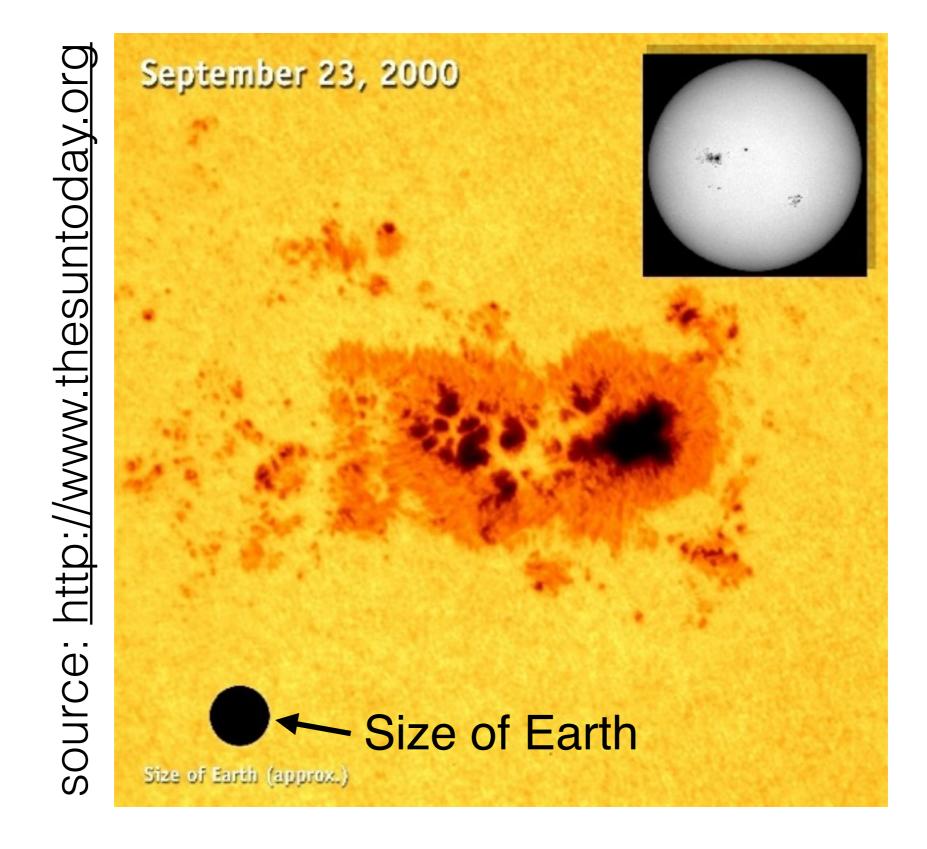


Why are those pictures of the same thing so different?

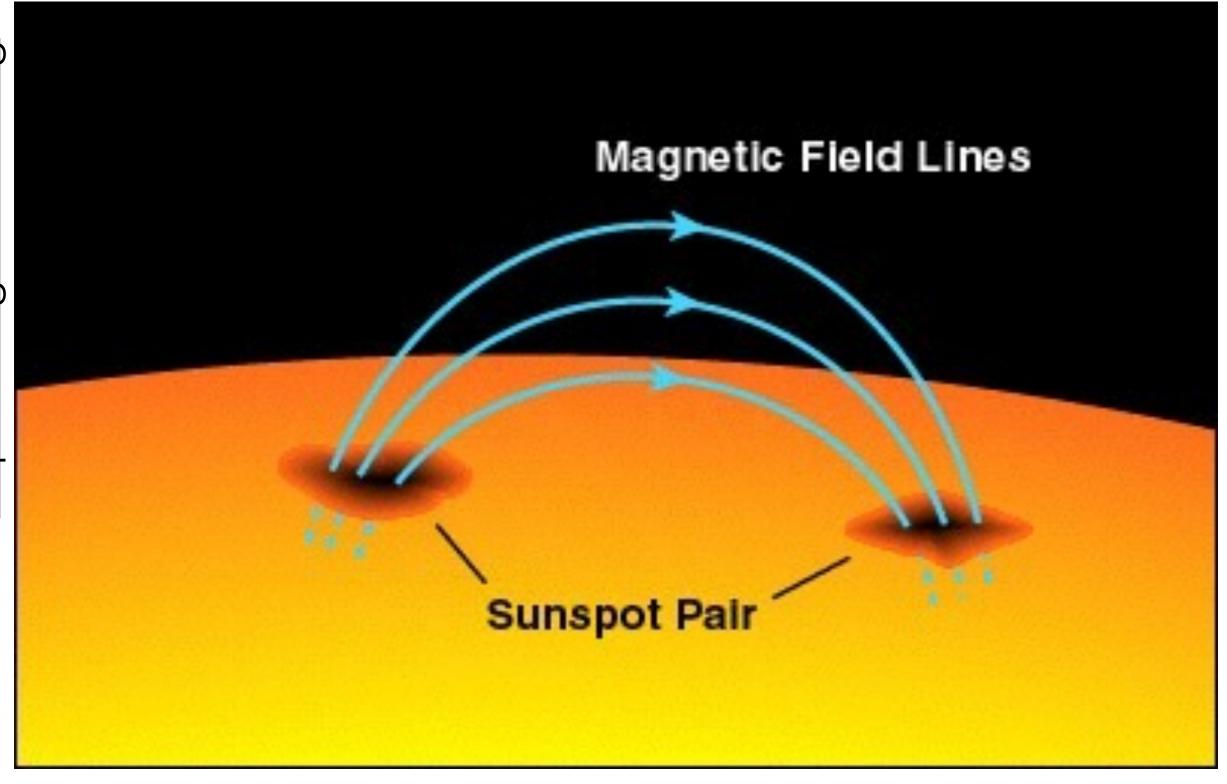




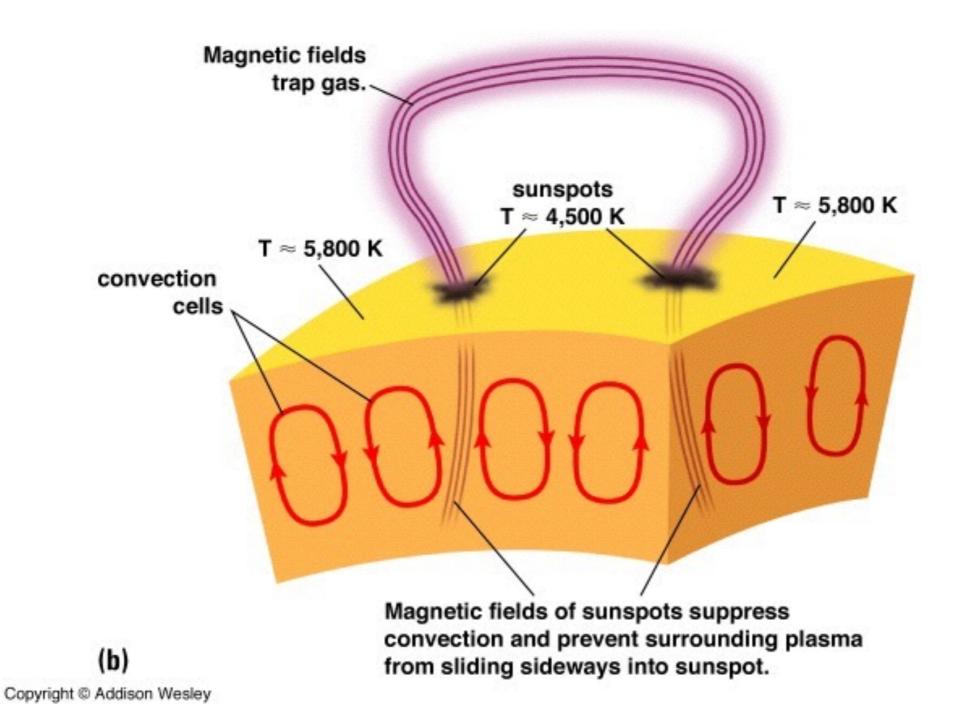
Sunspots form where concentrated magnetic field emerges through the photosphere.



Smaller sunspots are about the size of the Earth.

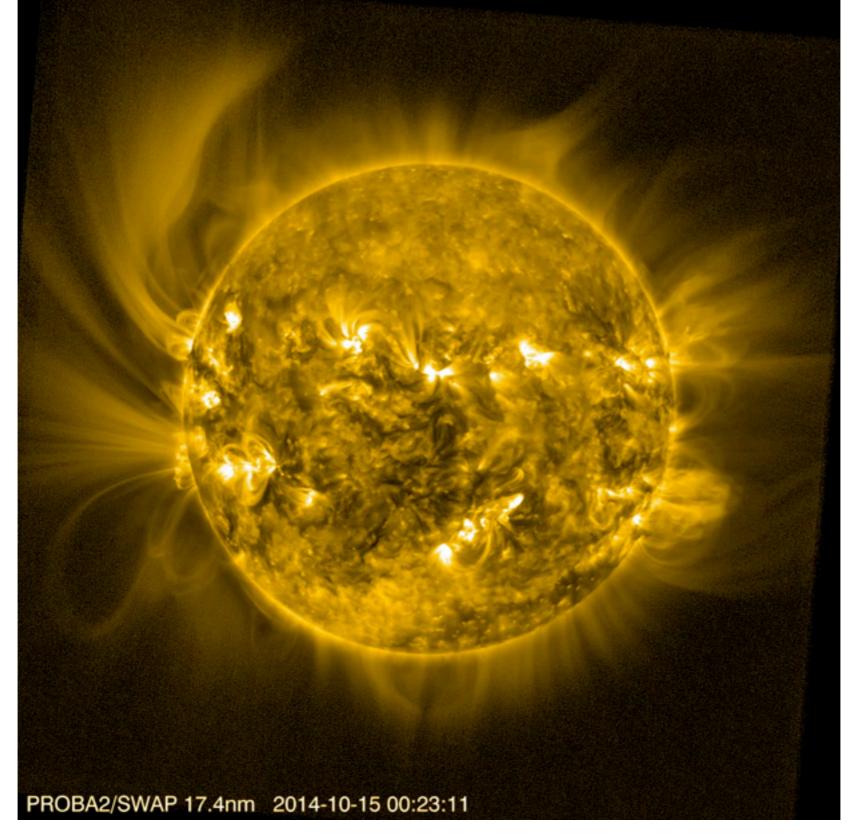


Strong magnetic field threads through sunspots.



Strong vertical field inhibits convection, making sunspots cooler than the surrounding photosphere.

Above Sunspots: Active Regions

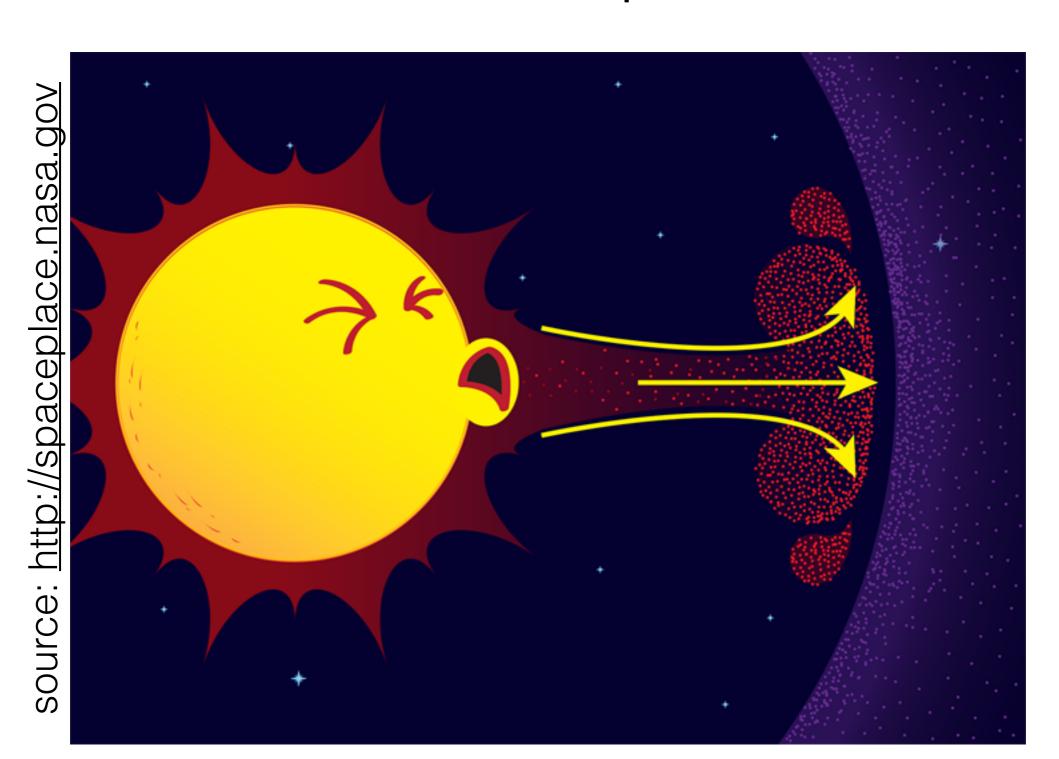


The dynamic corona

PROBA2/SWAP movie of 3 solar rotations

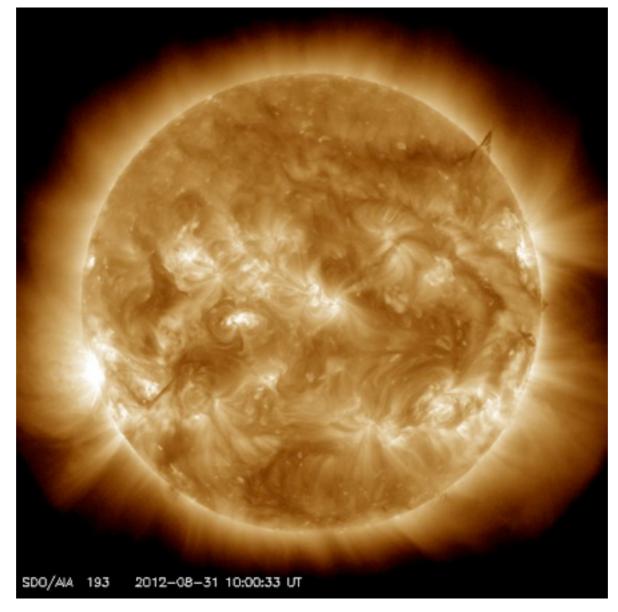
Heliosphere

The bubble-like volume surrounding solar system caused by the *solar wind*. Outside the heliosphere is *interstellar space*.



Eruptions

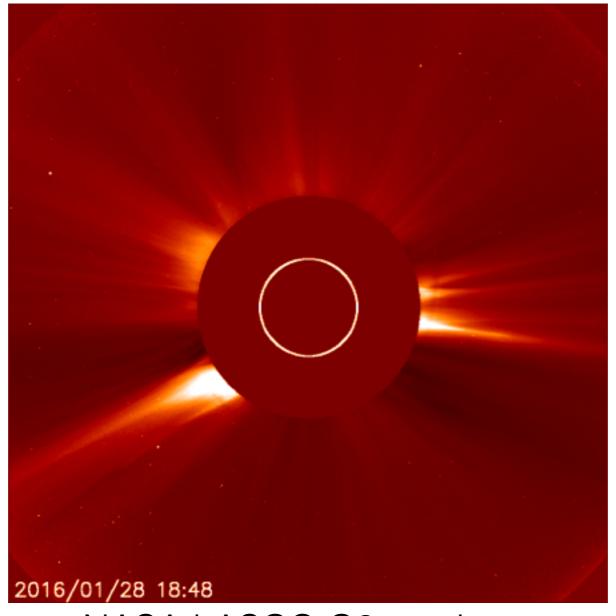
Major disturbances in the heliosphere are caused by massive explosions in the Sun's atmosphere: *coronal mass ejections*.



NASA SDO/AIA movie

Eruptions

Major disturbances in the heliosphere are caused by massive explosions in the Sun's atmosphere: *coronal mass ejections*.

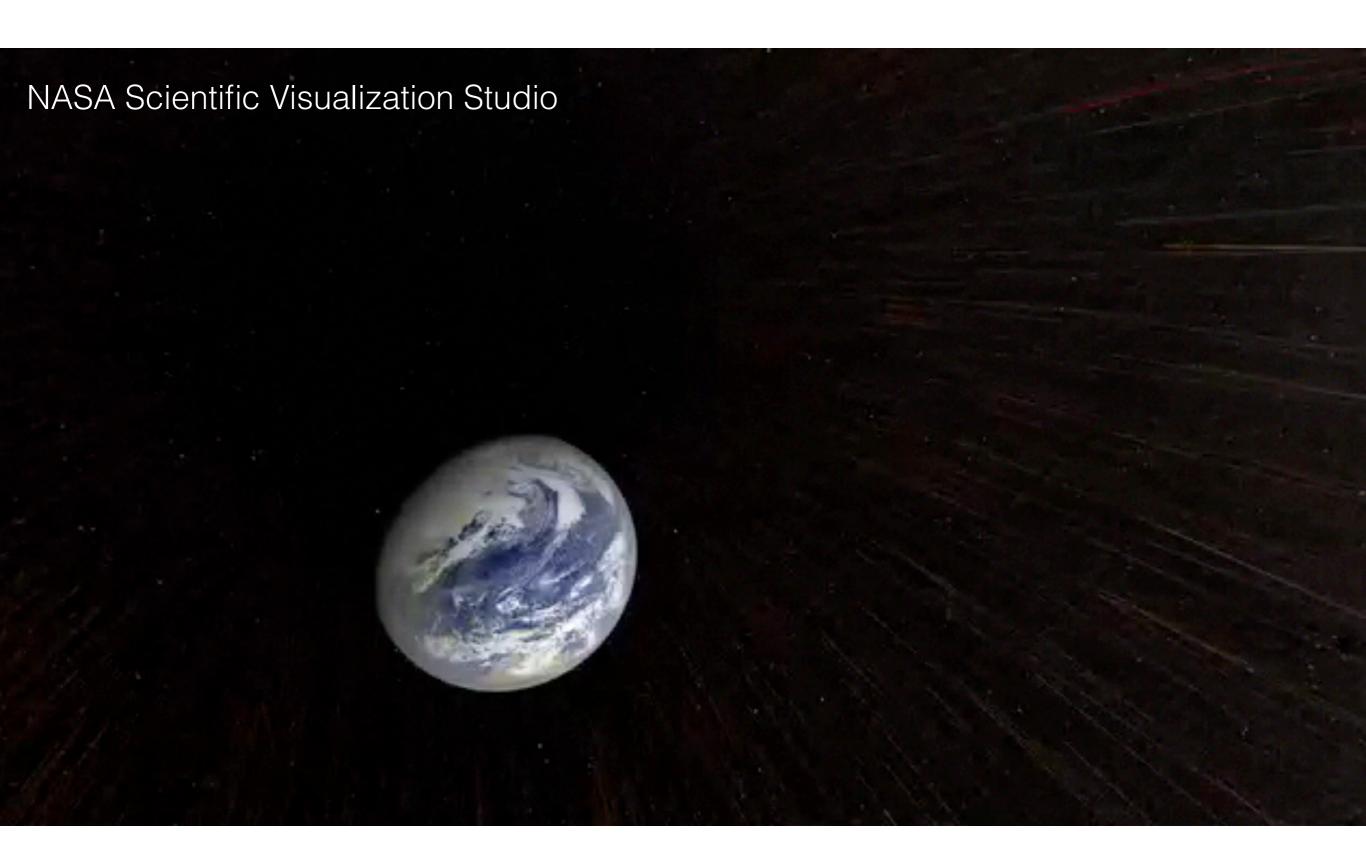


NASA LASCO C2 movie

Eruption statistics

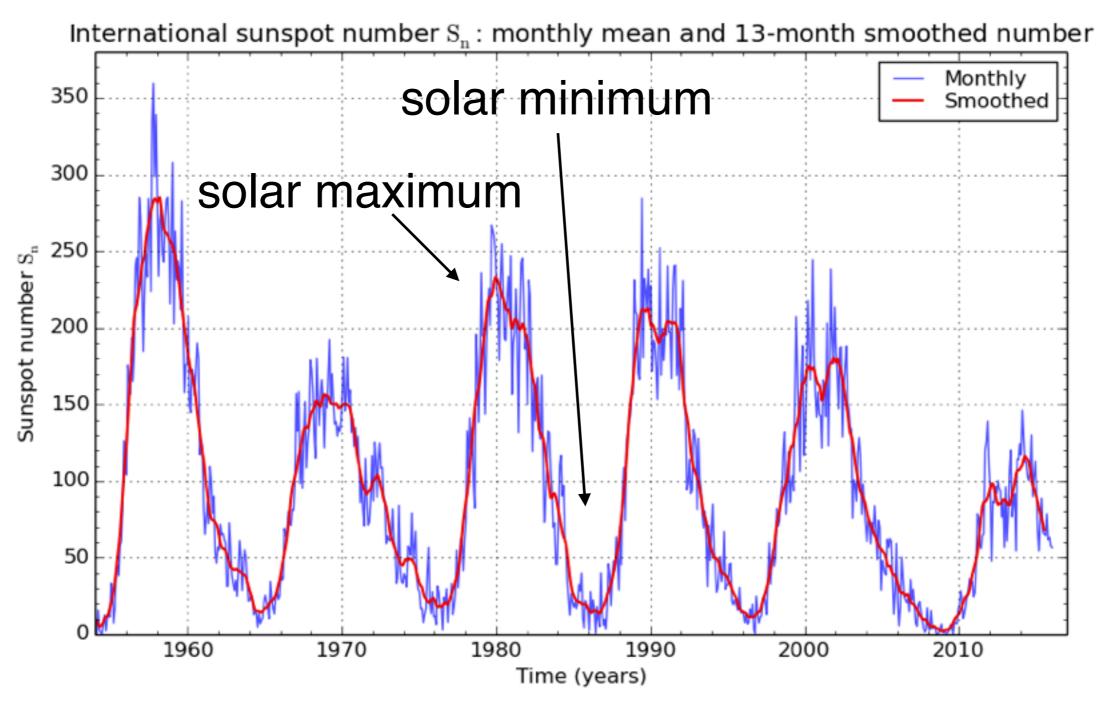
- How big? About as 1 cubic km³ of water
- How fast? About 500 km/s (1100 mph)
- How much energy? About 20x the last year's global energy consumption

Eruptions in the heliosphere

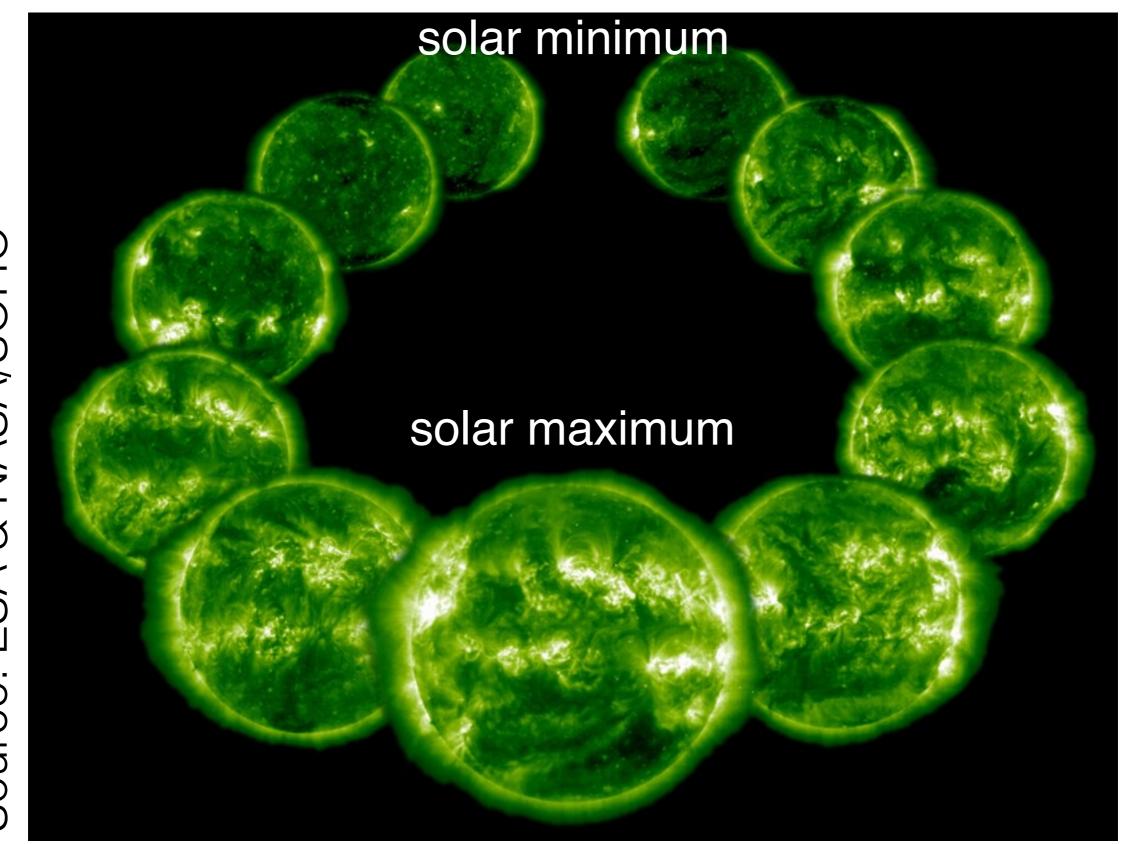


Activity cycle

There are times when the Sun is more active than others. It is linked to the solar dynamo. The activity cycle period is roughly 11 years.



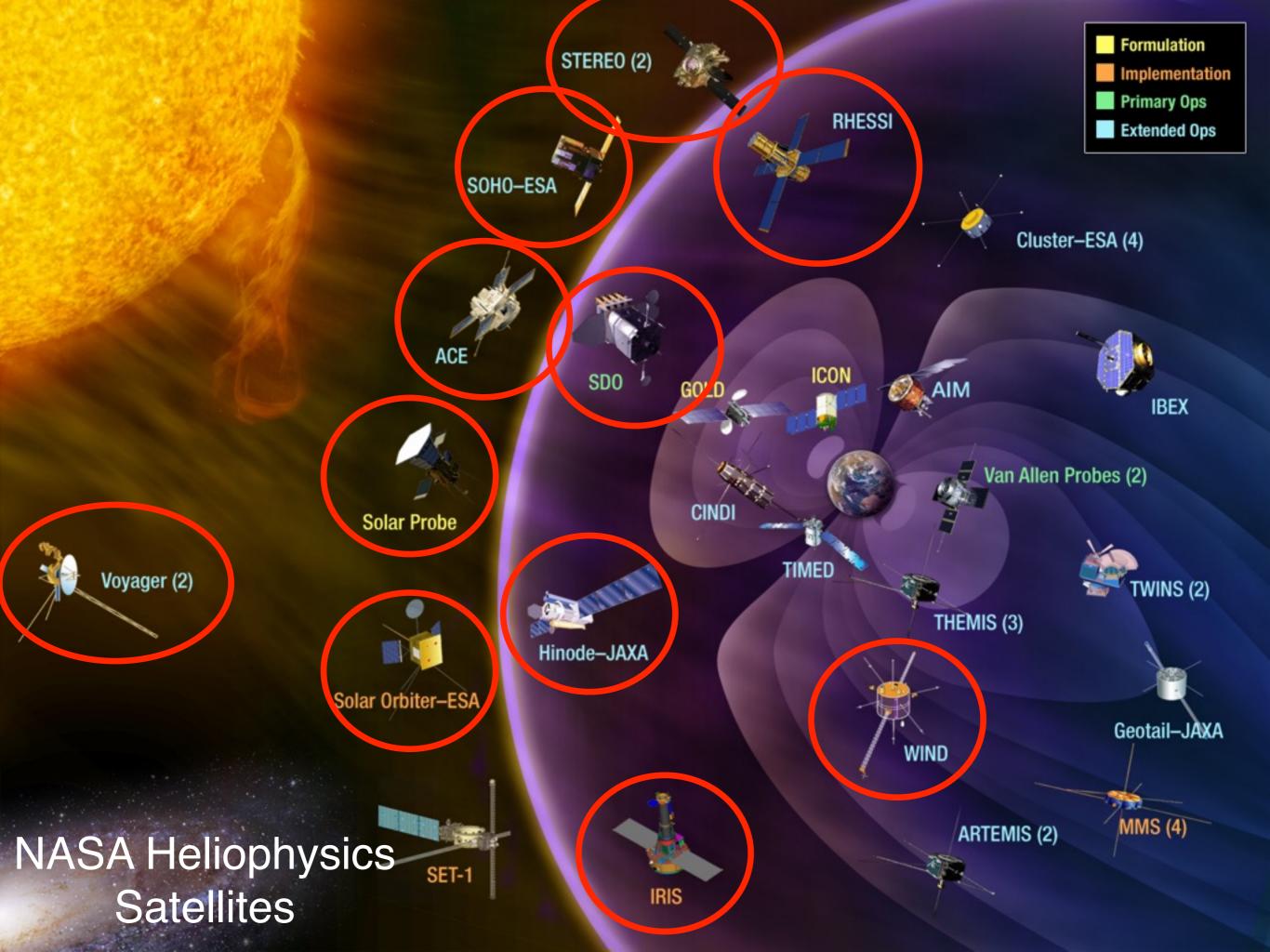
Activity cycle

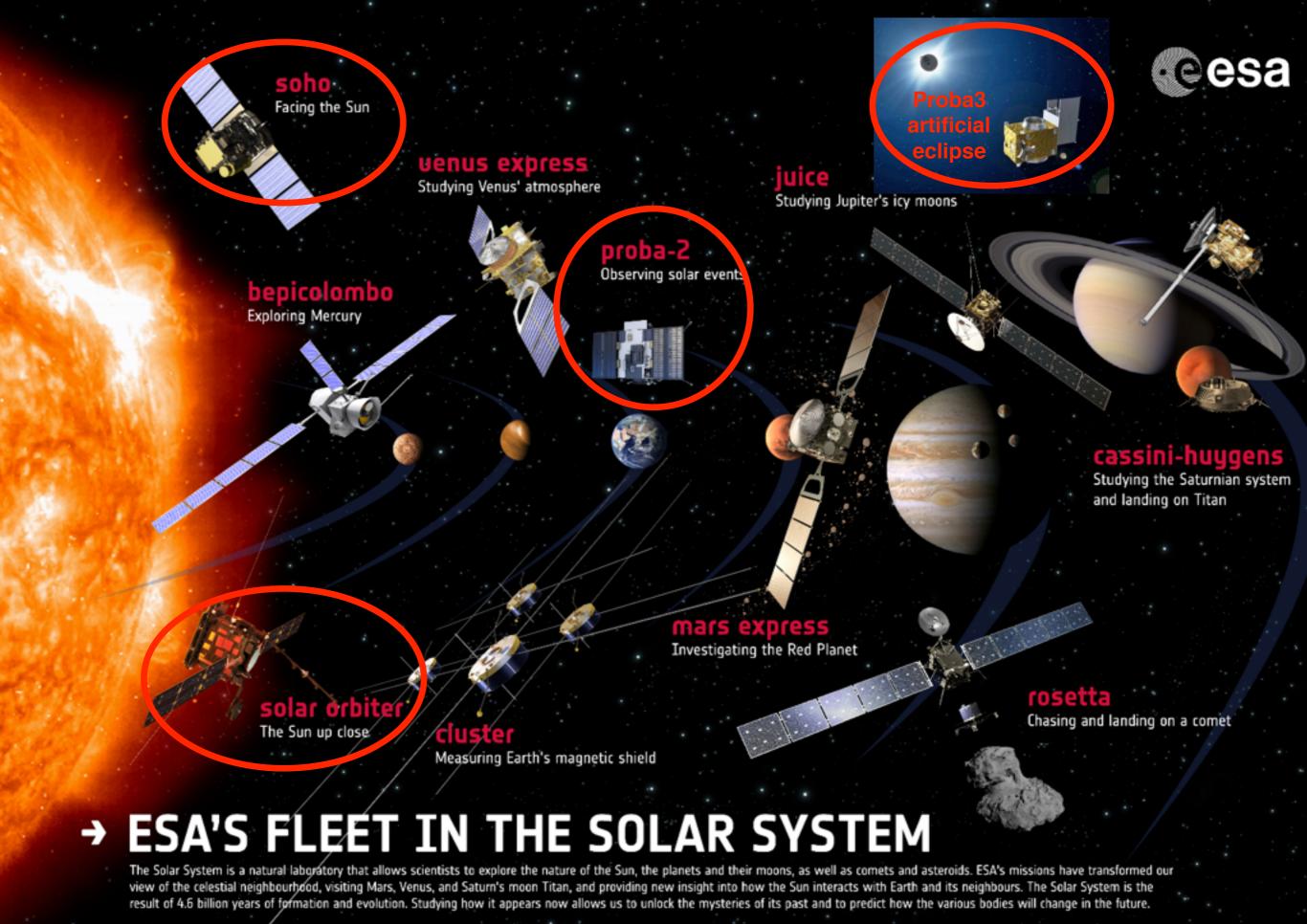


Source: ESA & NASA/SOHO

How do we know all of this?

- Solar data:
 - remote sensing: images, total brightness, spectra, polarimetry, helioseismology
 - in-situ plasma density, velocity, magnetic field information
- Computer modeling of the sun at all scales.



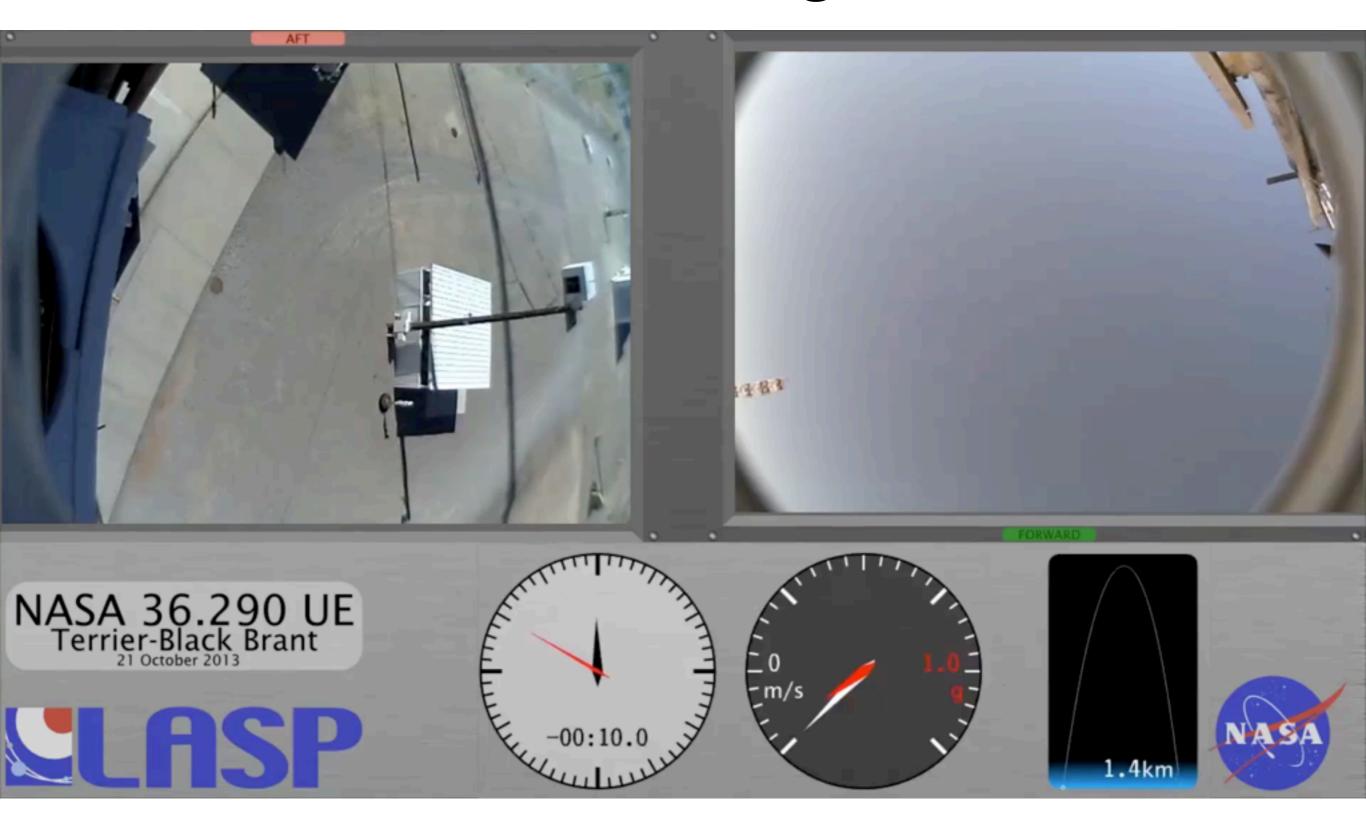


www.esa.int

CLASP launch 3 September 2015, White Sands Missile Range MSFC Sounding Rocket

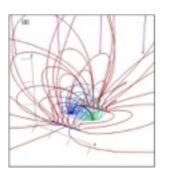


LASP Sounding Rocket



My research: magnetic fields in the corona

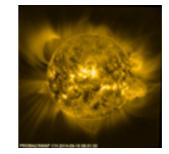
Computer models of eruptions.



Measurements of the coronal magnetic field.

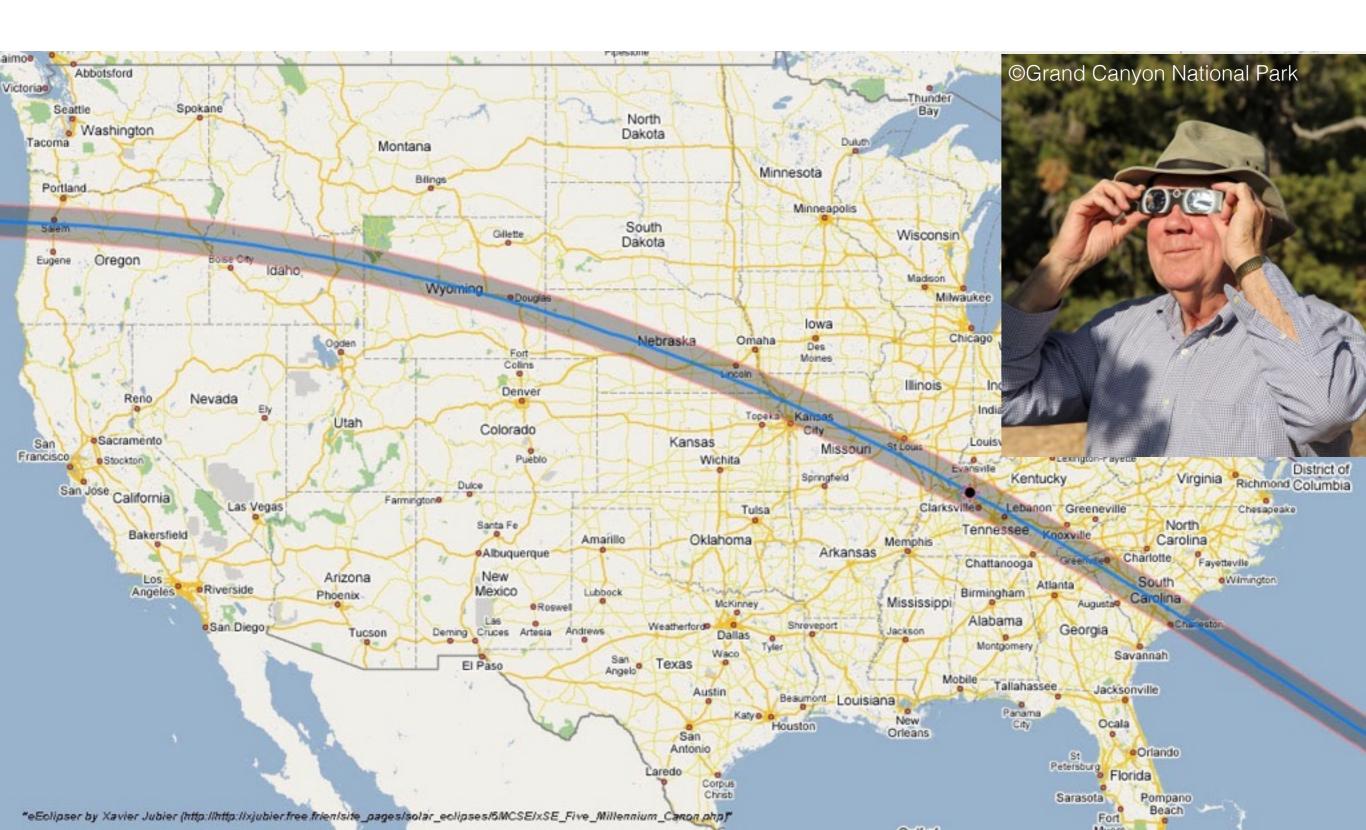


Large-scale structure of the corona.



 Sounding rockets, measurements of the magnetic field in the chromosphere.

Eclipse Aug 21, 2017!



Thank you!